

Computer Graphics

- Volume Rendering -

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Overview

- So far:
 - Light interactions with surfaces
 - Assume vacuum in and around objects
- This lecture:
 - Participating media
 - How to represent volumetric data
 - How to compute volumetric lighting effects
 - How to implement a very basic volume renderer

Fog & clouds

An aerial photograph of a dense forest during sunrise or sunset. The sun is low on the horizon, creating a strong glow and casting long, golden rays of light through the trees. The ground is covered in a thick layer of fog or mist, which catches the light and creates a hazy, ethereal atmosphere. The trees are mostly tall and thin, with their tops silhouetted against the bright sky. The overall color palette is dominated by warm, golden-brown and green tones.

Underwater



source: dailypictures.info

Surface or volume?



source: Flickr

source: Studio Lernert & Sander





Avatar. Copyright © 2009 20th Century Fox



Arrival. Copyright © 2016 Paramount Pictures



Big Hero 6. Copyright © 2014 Walt Disney Enterprises, Inc.



Mortal Engines. Copyright © 2018 Universal Studios

Fundamentals

Volumetric Effects

- Light interacts not only with surfaces but everywhere inside!
- Volumes scatter, emit, or absorb light



<http://coclouds.com>



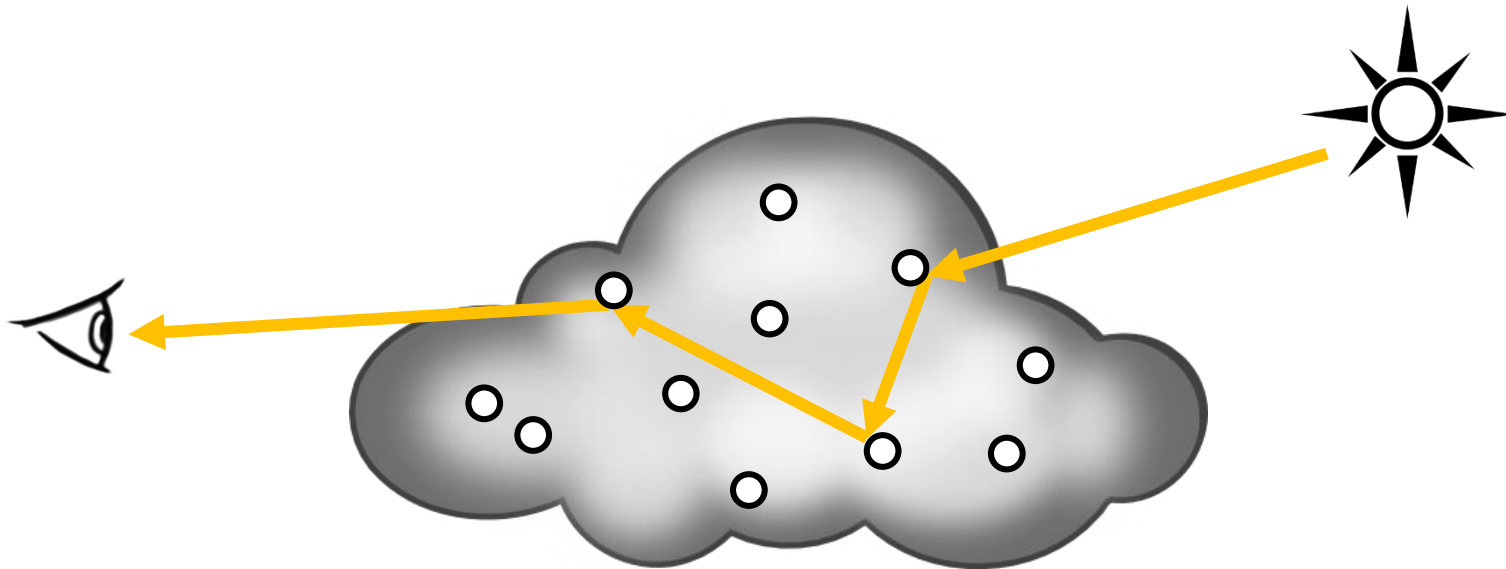
<http://wikipedia.org>



<http://commons.wikimedia.org>

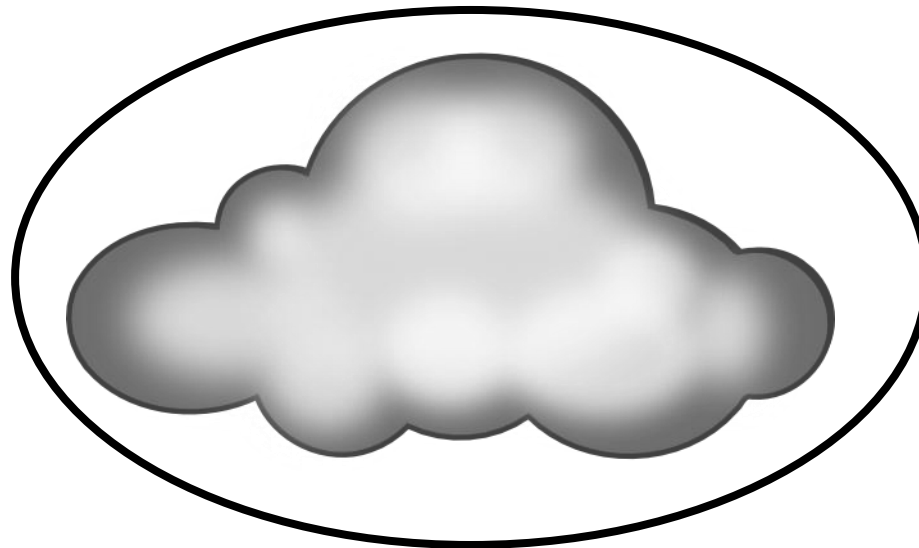
Approximation: Model Particle Density

- Modeling individual particles of a volume is, of course, not practical
- Instead, represent statistically using the average density
- (Same idea as, e.g., microfacet BSDFs)



Volume Representation

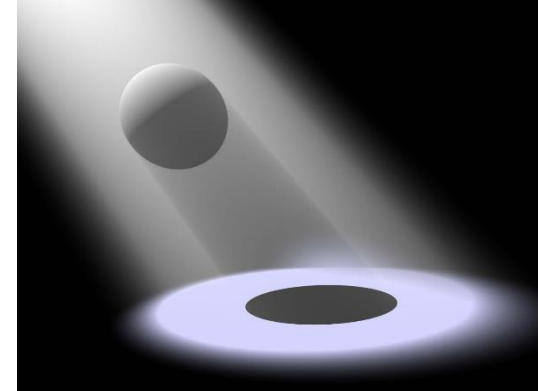
- Many possibilities (particles, voxel octrees, procedural,...)
- A common approach: Scene objects can “contain” a volume



Volume Representation

- Homogeneous:
 - Constant density
 - Constant absorption, scattering, emission,
 - Constant phase function (later)

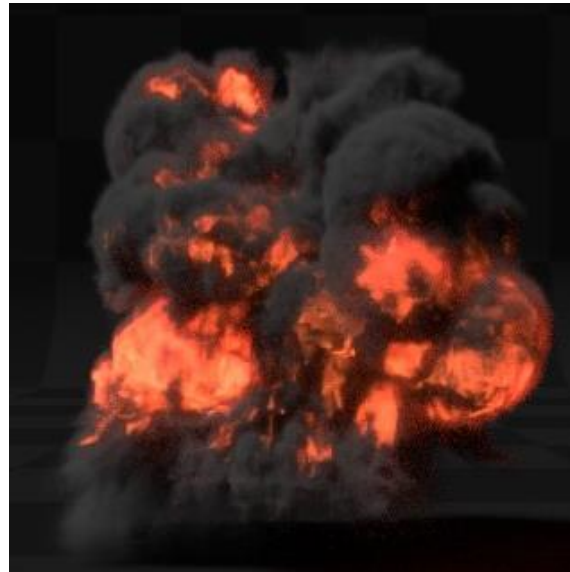
- Heterogeneous:
 - Coefficients and/or phase function vary across the volume
 - Can be represented using **3D textures**
 - (e.g., voxel grid, procedural)



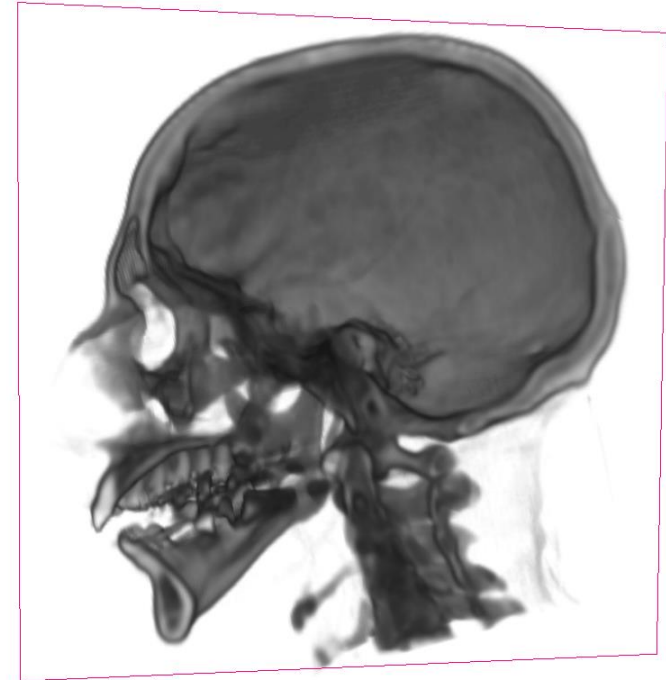
<http://wikipedia.org>

Data Acquisition

- Real-world measurements via tomography
- Simulation, e.g.,
 - Fluids,
 - Fire and smoke,
 - Fog



<https://docs.blender.org>



Simulating Volumes

Mathematical Formulation of Volumetric Light Transport

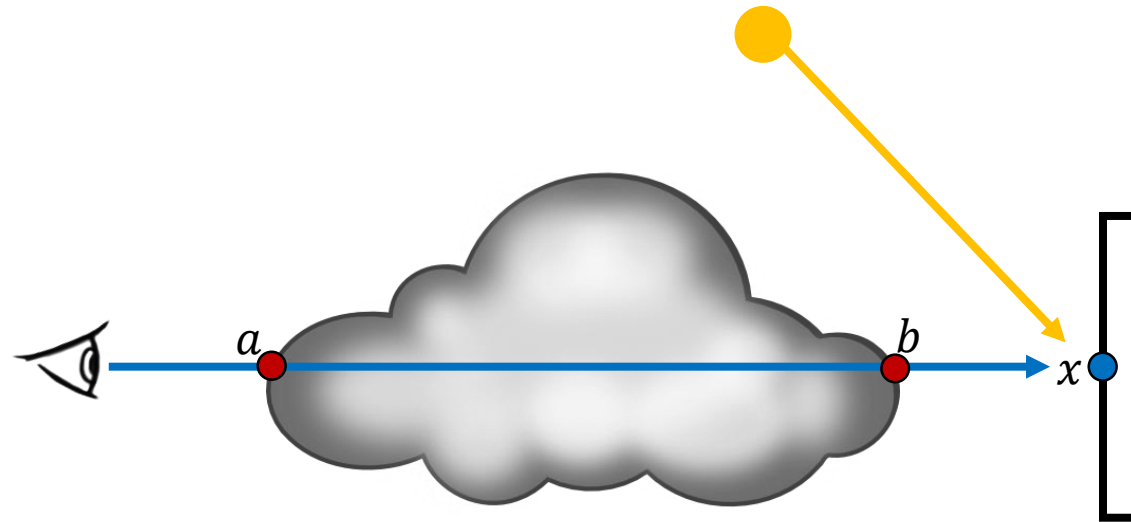
So far: Assume Vacuum

- Compute $L_o(x, \omega_o)$ using the rendering equation



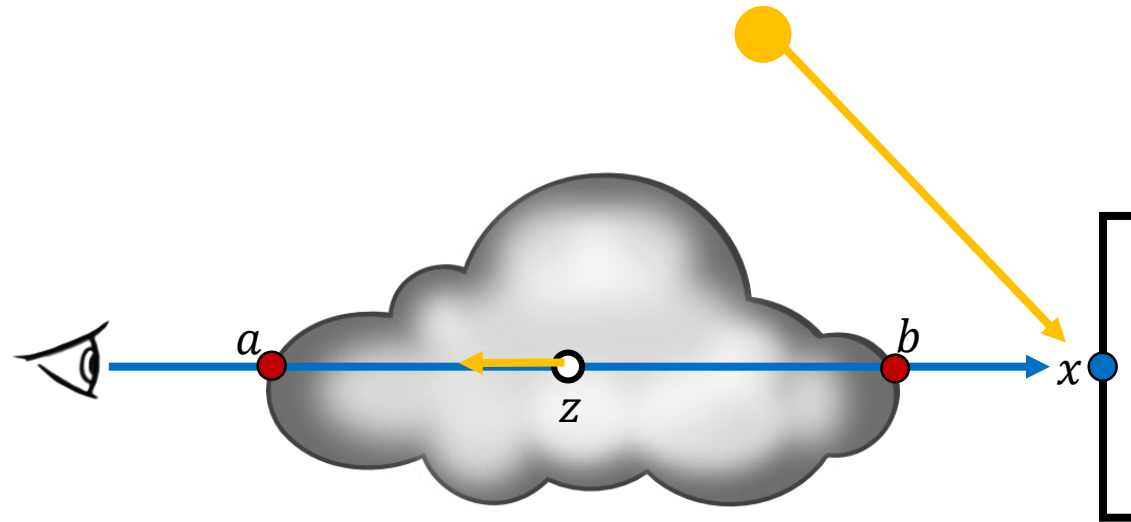
Volume Absorbs and Scatters Light

- Compute $L_o(x, \omega_o)$ using the rendering equation
- Only a fraction $T(a, b)L_o(x, \omega_o)$ arrives at the eye



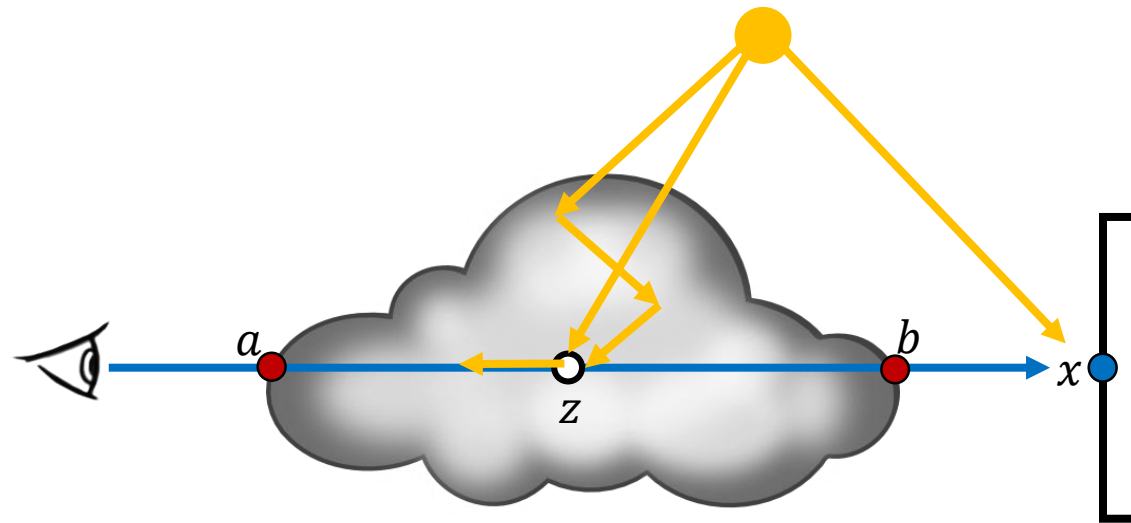
Volume Emits Light

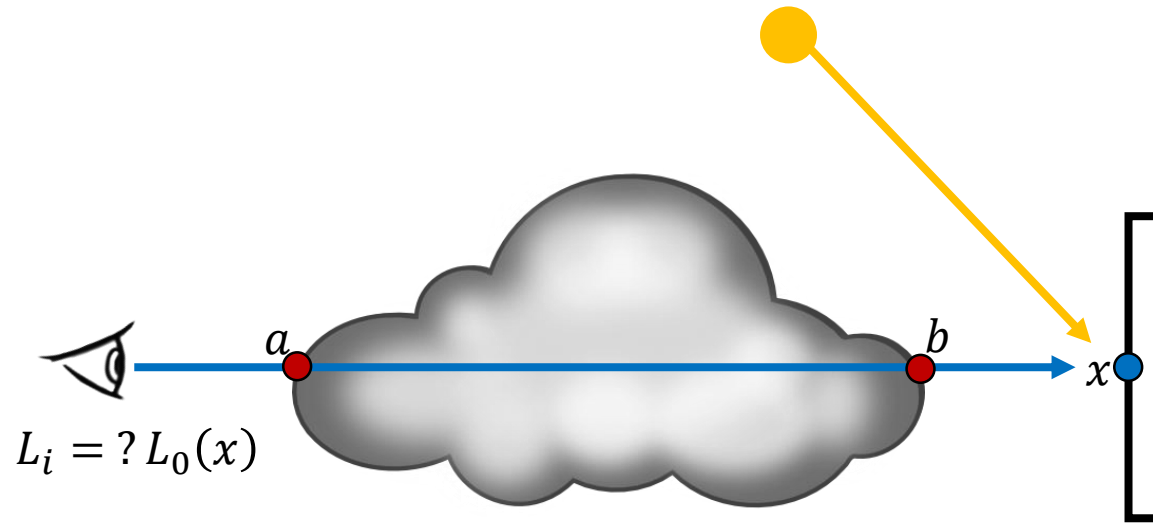
- Compute $L_o(x, \omega_o)$ using the rendering equation
- Only a fraction $T(a, b)L_o(x, \omega_o)$ arrives at the eye
- Every point z between a and b might emit light



Volume Scatters Light

- Compute $L_o(x, \omega_o)$ using the rendering equation
- Only a fraction $T(a, b)L_o(x, \omega_o)$ arrives at the eye
- Every point z between a and b might emit light
- Every point z might be illuminated through the volume





Attenuation

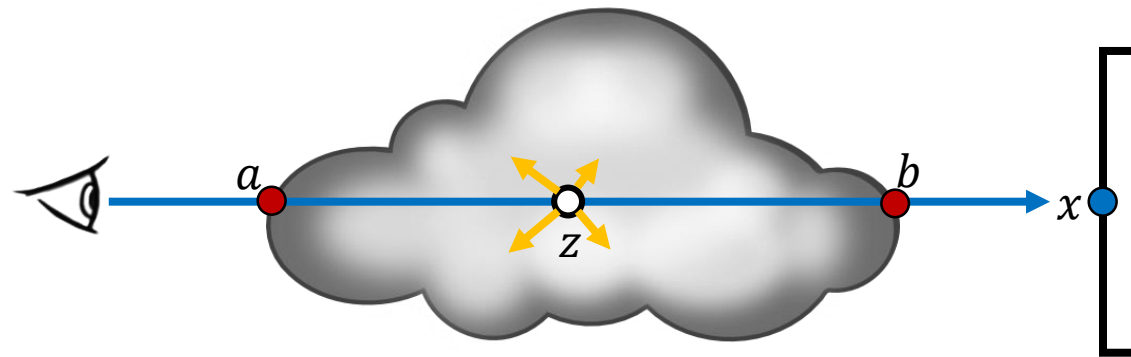
Computing Absorption and Out-Scattering



<http://commons.wikimedia.org>

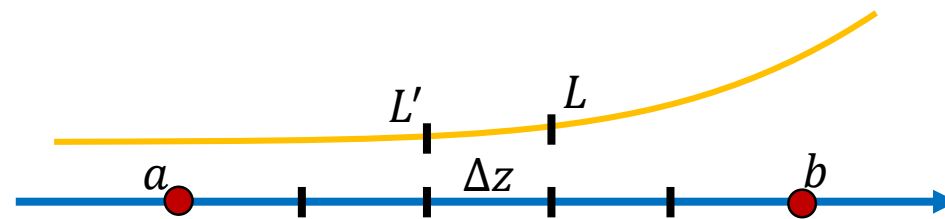
Attenuation = Absorption + Out-Scattering

- Every point in the volume might absorb light or scatter it in other directions
- Modeled by absorption and scattering densities: $\mu_a(z)$ and $\mu_s(z)$
- Might depend on position, direction, time, wavelength,...
- For simplicity: we assume only positional dependence



Computing Absorption – Intuition

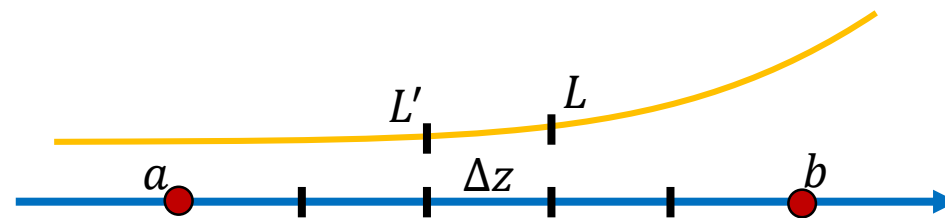
- Consider a small segment Δz
- Along that segment, radiance is reduced from L to L'
- $L' = L - \mu_a L \Delta z$
- Where μ_a is the percentage of radiance that is absorbed (per unit distance)



Computing Absorption – Intuition

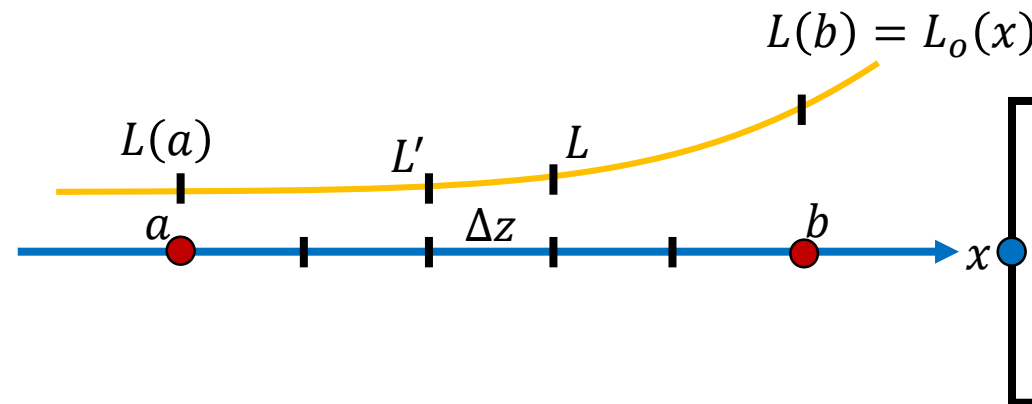
- Consider a small segment Δz
- Along that segment, radiance is reduced from L to L'
- $L' = L - \mu_a L \Delta z$
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- Lets rewrite this:
- $\Delta L = L' - L = -\mu_a L \Delta z$



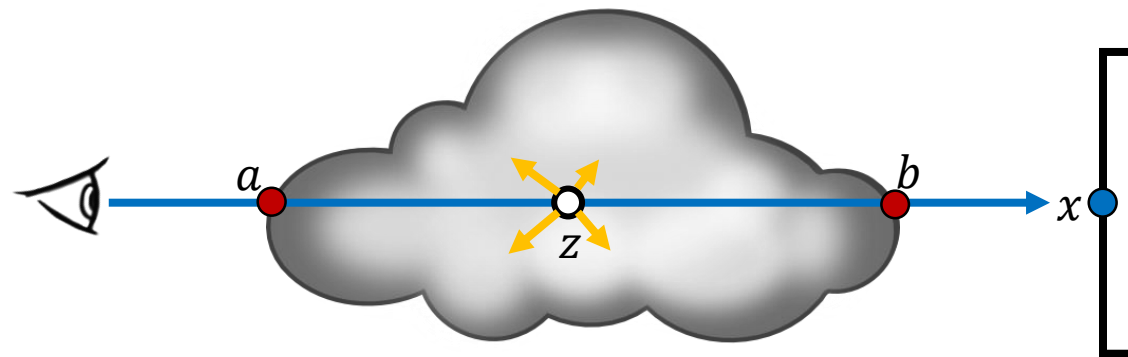
Computing Absorption – Exponential Decay

- $\Delta L = -\mu_a L \Delta z$
- For infinitely small Δz , this becomes
- $dL = -\mu_a L dz$
- A differential equation that models exponential decay!
- Solution: $L(a) = L_o(x) e^{-\int_b^a \mu_a(t) dt}$



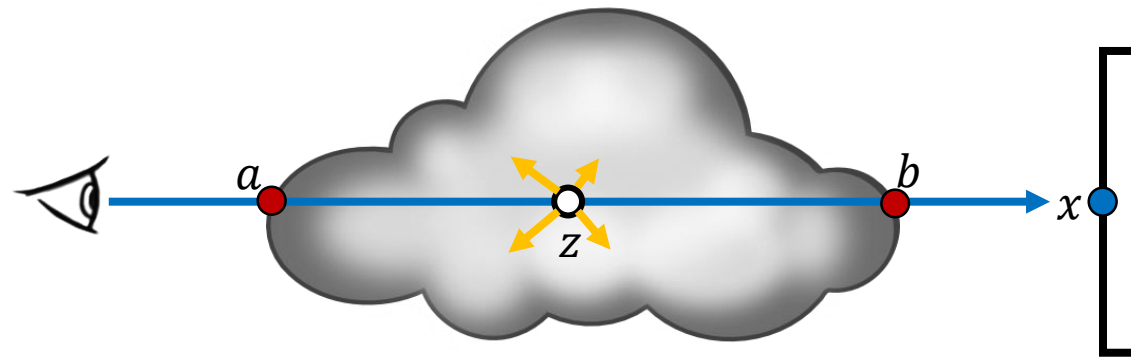
Computing Out-Scattering

- Same as absorption, only different factor!
- $\mu_s(z)$: percentage of light scattered at point z
- $L(a) = L_o(x) e^{-\int_b^a \mu_s(t) dt}$



Computing Attenuation

- Fraction of light that is neither absorbed nor out-scattered
- $\mu_t = \mu_a + \mu_s$
- $L(a) = L_o(x) e^{-\int_b^a (\mu_a(t) + \mu_s(t)) dt}$
- Attenuation: $T(a, b) = e^{-\int_b^a \mu_t(t) dt}$

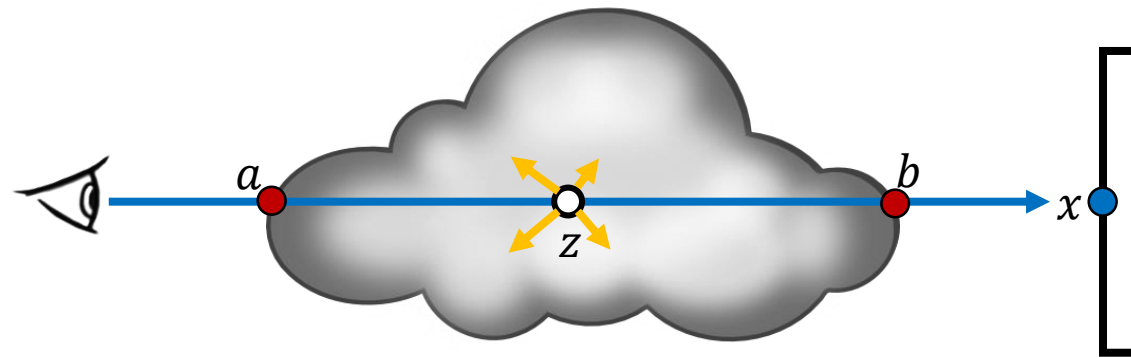


Computing Attenuation – Homogeneous

- Simple case: constant density / attenuation

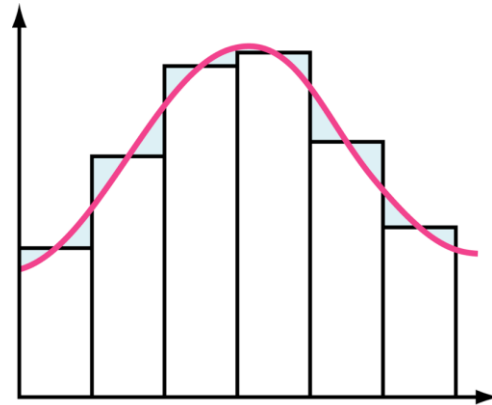
- $\mu_t(z) = \mu_t \quad \forall z$

- $T(a, b) = e^{-\int_b^a \mu_t(t) dt} = e^{-(a-b)\mu_t}$



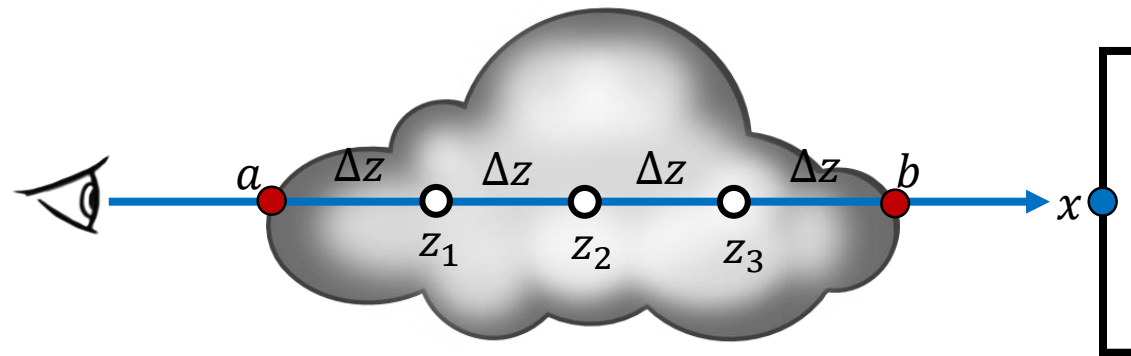
Estimating Attenuation

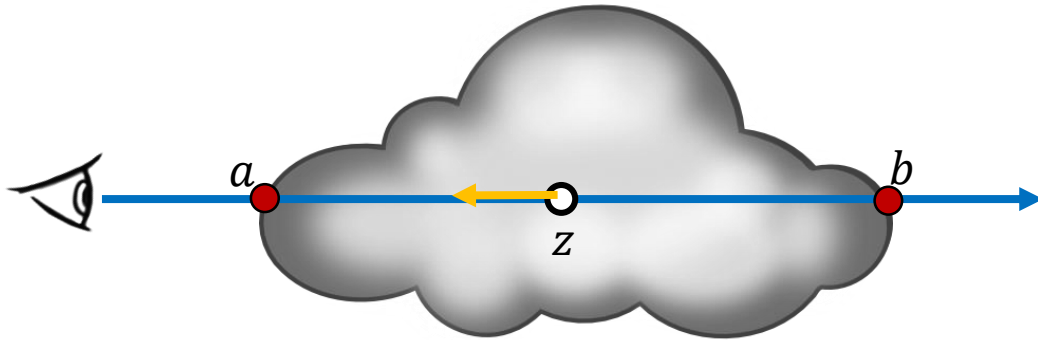
- We need to solve another integral:
 - $T(a, b) = e^{-\int_b^a \mu_t(t) dt}$
- Many solutions, e.g., Monte Carlo integration (next semester)
- Simple solution: Quadrature



Estimating Attenuation – Ray Marching

- We need to solve another integral:
 - $T(a, b) = e^{-\int_b^a \mu_t(t) dt}$
- Simple solution: Quadrature
- Ray marching: evaluate at discrete positions (fixed stepsize Δz)
- $\int_b^a \mu_t(t) dt \approx \sum_i \mu_t(z_i) \Delta z$





Emission

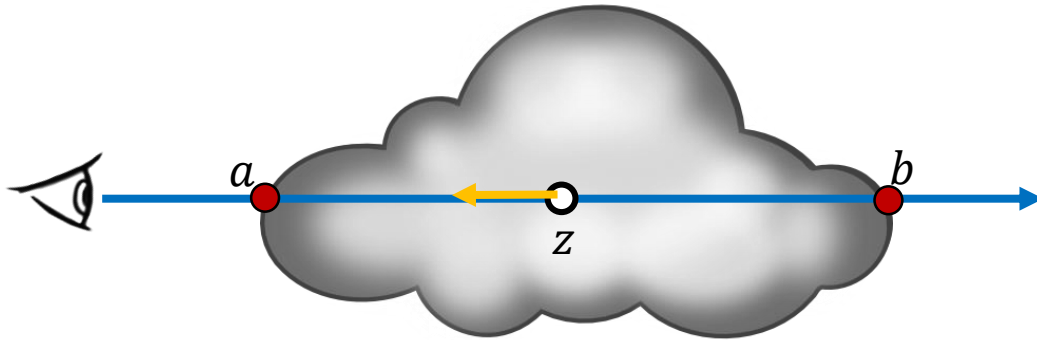
Explosions!



<http://wikipedia.org>

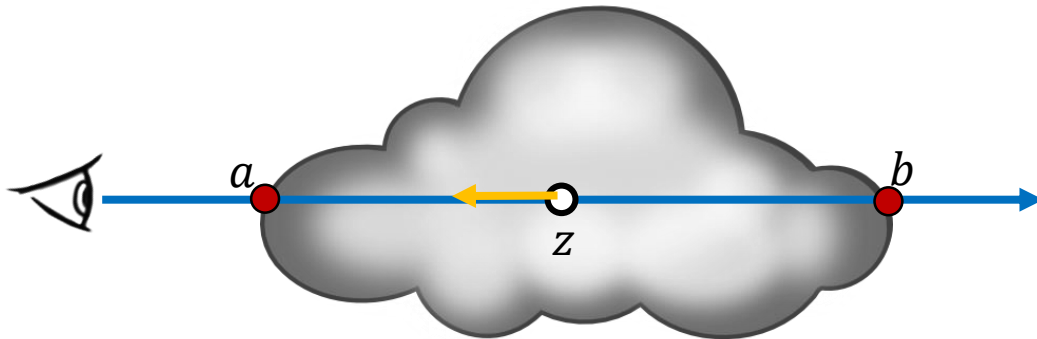
Every Point Might Emit Light

- Assume z emits $L_e(z)$ towards a
- Some of that light might be absorbed or out-scattered: It is attenuated
- $L(a) = L_e(z) T(z, a)$



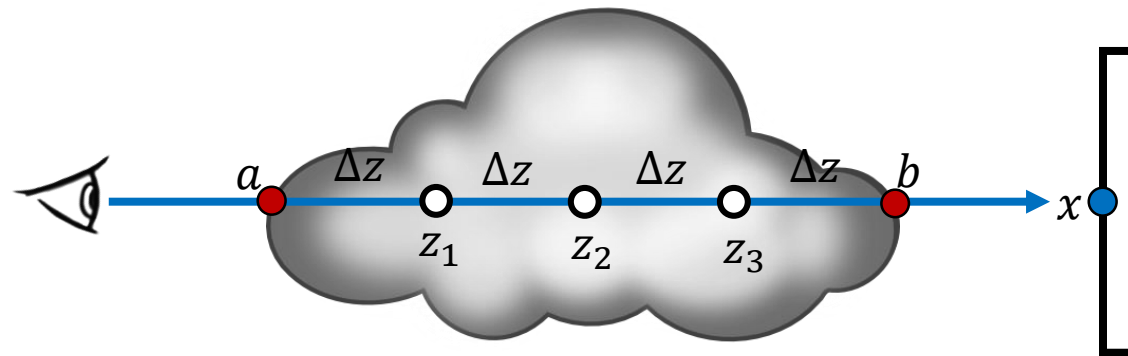
Every Point Might Emit Light

- Assume z emits $L_e(z)$ towards a
- Some of that light might be absorbed or out-scattered: It is attenuated
- $L(a) = L_e(z) T(z, a)$
- Happens at every point along the ray!
- $L(a) = \int_a^b L_e(z) T(z, a) dz$
- Another integral...



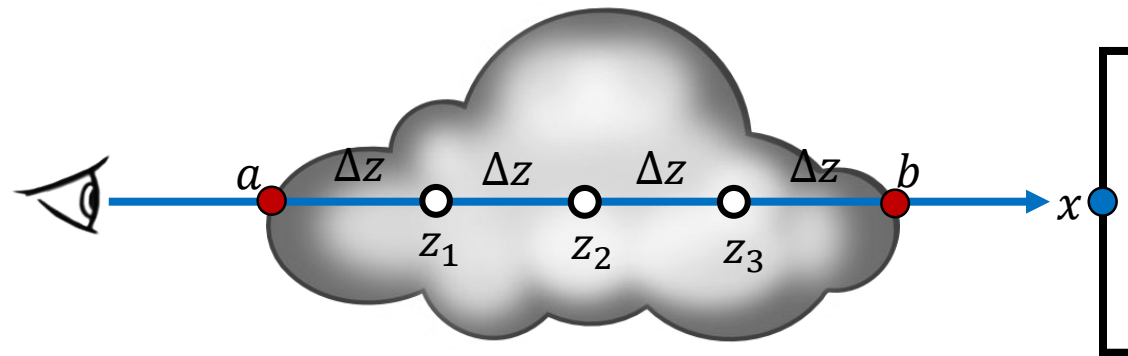
Ray Marching for Emission

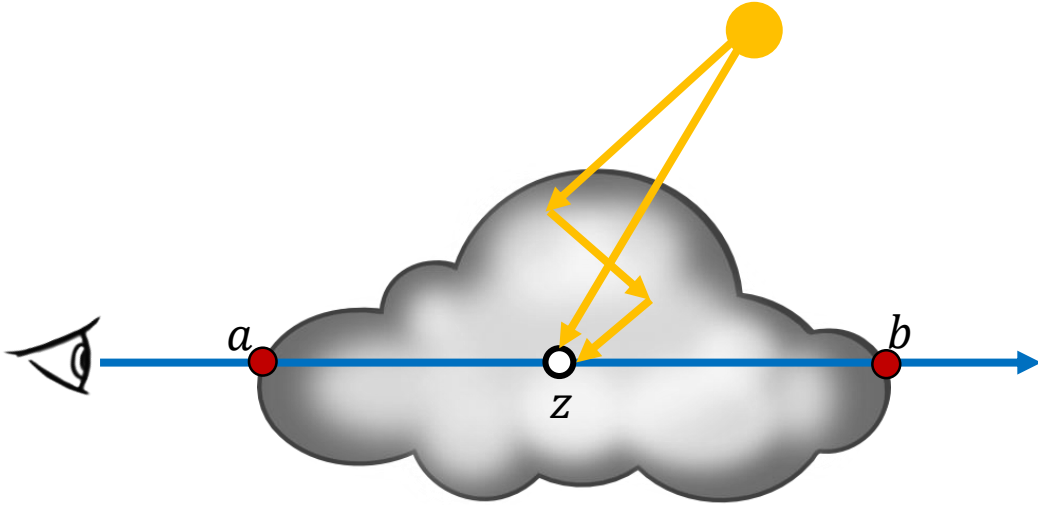
- Same as before: integrate via quadrature
- $\int_a^b L_e(z) T(z, a) dz \approx \sum_i L_e(z_i) T(z_i, a) \Delta z$
- Attenuation $T(z_i, a)$ estimated as before



Ray Marching for Emission

- Same as before: integrate via quadrature
- $\int_a^b L_e(z) T(z, a) dz \approx \sum_i L_e(z_i) T(z_i, a) \Delta z$
- Attenuation $T(z_i, a)$ estimated as before
- Attenuation can be incrementally updated:
 - $T(z_i, a) = T(z_{i-1}, a) T(z_i, z_{i-1})$
 - (because it is an exponential function)





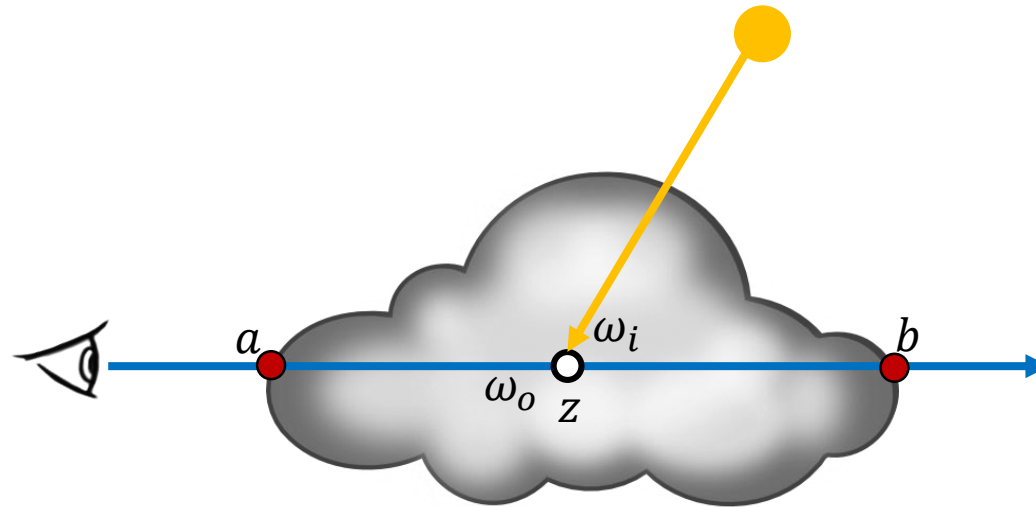
In-Scattering

Accounting for Reflections Inside the Volume



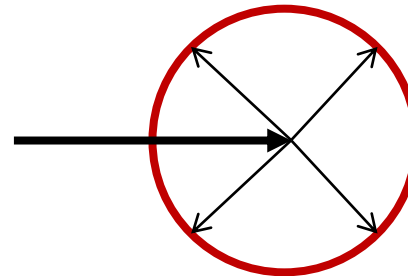
Direct Illumination

- Account for the (attenuated) direct illumination at every point z
- Similar to the rendering equation:
- $L_o(z, \omega_o) = \int_{\Omega} L_i(x, \omega_i) f_p(\omega_i, \omega_o) d\omega_i$
- Integration over the whole sphere Ω
- The **phase function** f_p takes on the role of the BSDF



Phase Functions

- $L_o(z, \omega_o) = \int_{\Omega} L_i(x, \omega_i) f_p(\omega_i, \omega_o) d\omega_i$
- Describe what fraction of light is reflected from ω_i to ω_o
- Similar to BSDF for surface scattering
- Simplest example: isotropic phase function
 - $f_p(\omega_i, \omega_o) = \frac{1}{4\pi}$
 - (energy conservation: $\int_{\Omega} \frac{1}{4\pi} d\omega = 1$)



Phase Functions: Henyey-Greenstein

- Widely used
- Easy to fit to measured data
- $f_p(\omega_i, \omega_o) = \frac{1}{4\pi} \frac{1-g^2}{(1+g^2+2g \cos(\omega_i, \omega_o))^{\frac{3}{2}}}$
- g : asymmetry (scalar)
- $\cos(\omega_i, \omega_o)$: cosine of the angle formed by ω_i and ω_o

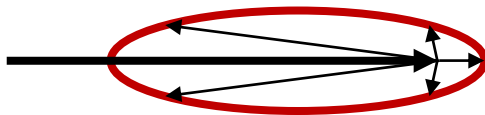
Henyey-Greenstein: Asymmetry Parameter

- $g = 0$: isotropic
- Negative g : back scattering
- Positive g : forward scattering

Back Scattering



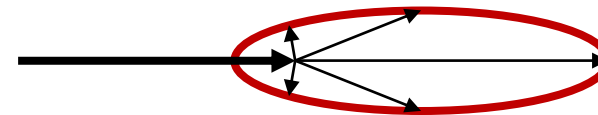
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Forward Scattering

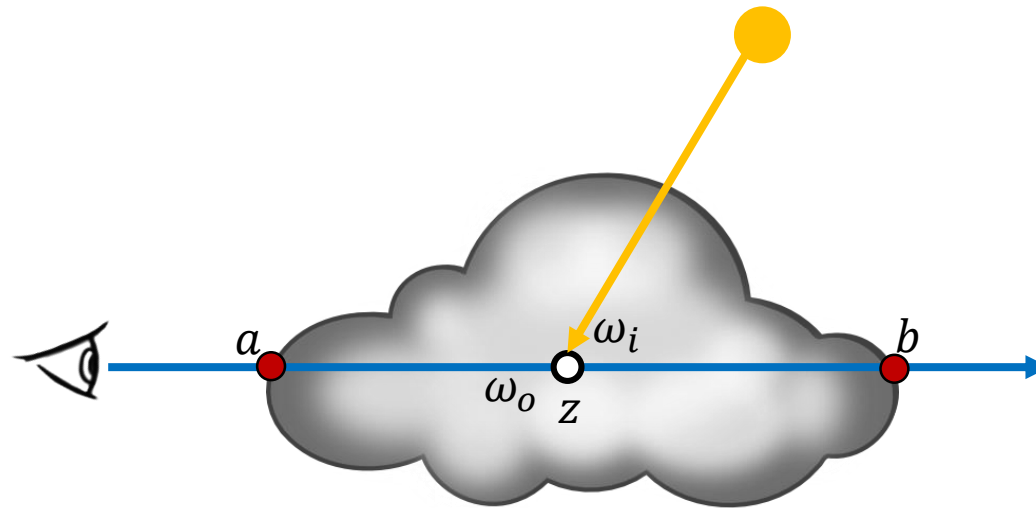


<http://coclouds.com>



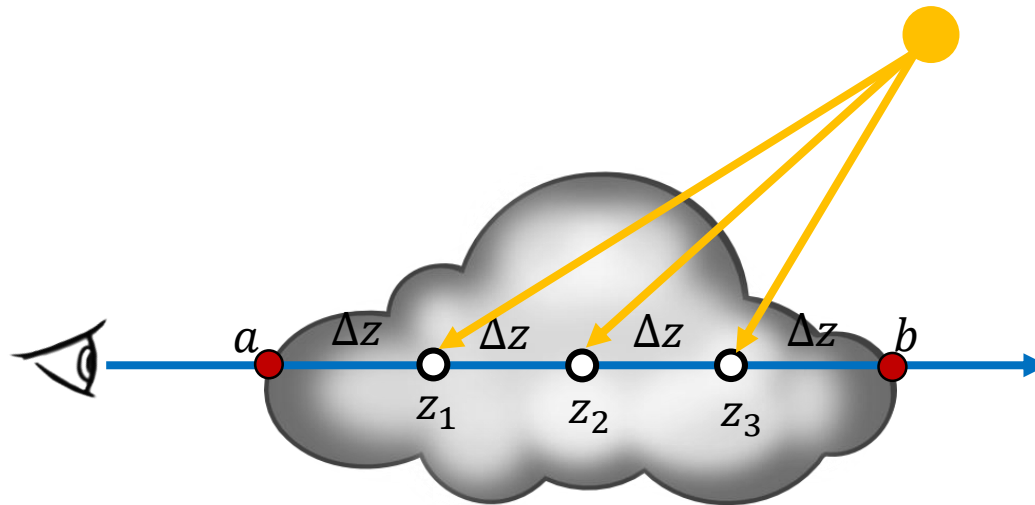
How to Estimate Volume Direct Illumination

- Reflected radiance at a point z : $L_o(z, \omega_o) = \int_{\Omega} L_i(x, \omega_i) f_p(\omega_i, \omega_o) d\omega_i$
- In our framework:
 - Sum over all point lights (as for surfaces)
 - Trace shadow ray (as for surfaces)
 - Estimate attenuation along the shadow ray (as for surfaces)



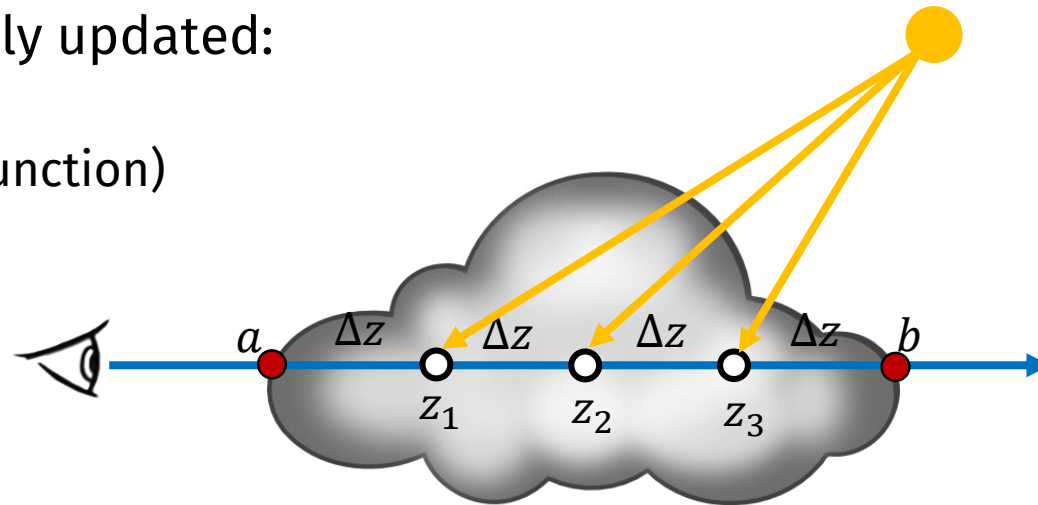
Ray Marching to Compute In-Scattering

- Same as for emission
- Goal: estimate the integral $\int_a^b T(z, a) \mu_s(z) L_i(z) f_p dz$
- Quadrature:
 - $\int_a^b T(z, a) \mu_s(z) L_i(z) f_p dz \approx \sum_i T(z_i, a) \mu_s(z) L_i(z_i) f_p \Delta z$



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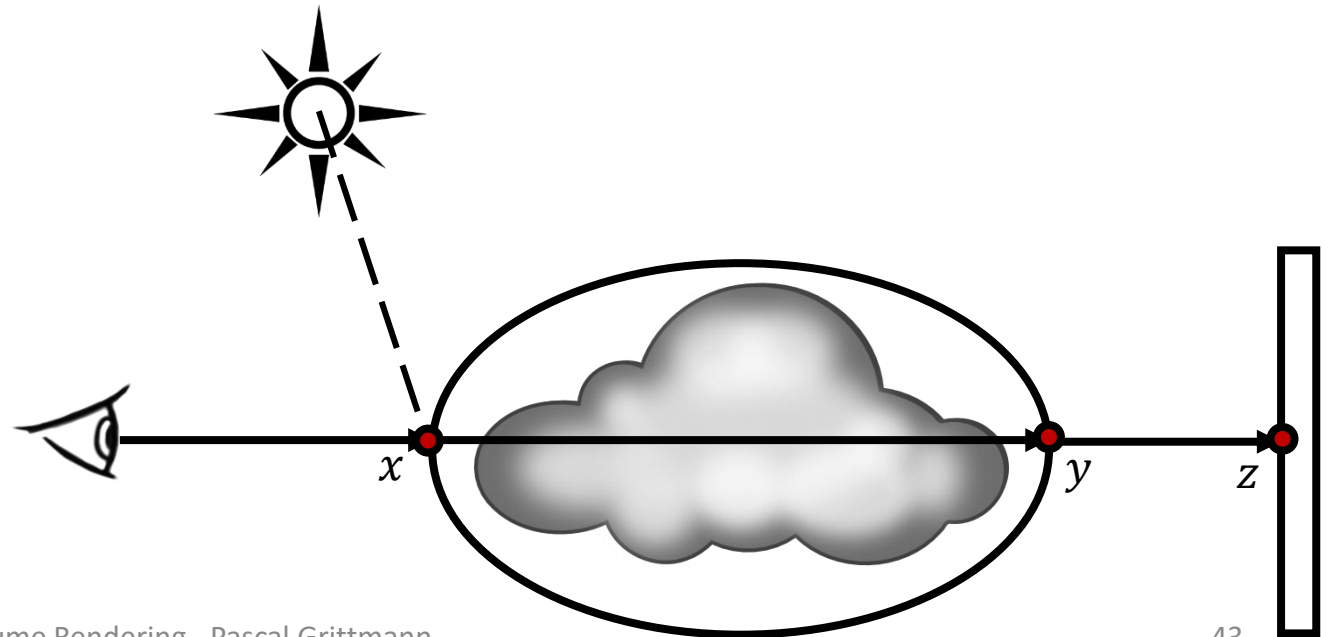


Putting it all Together

A Simple Volume Integrator

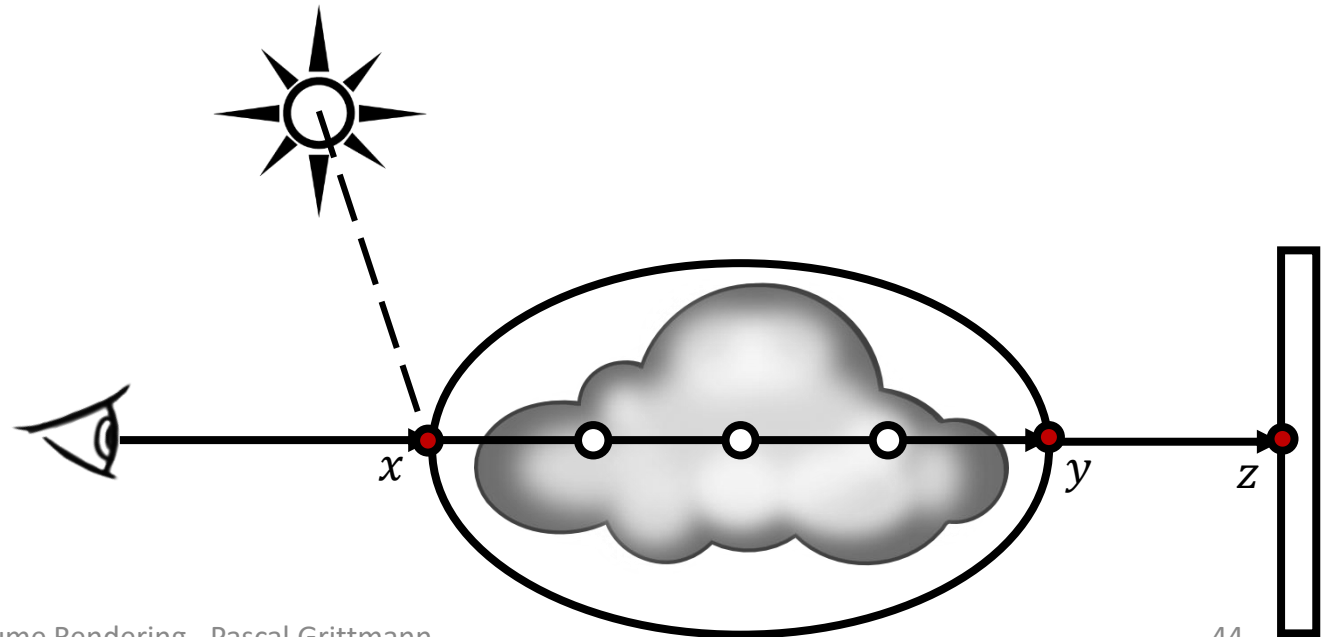
A Simple Volume Integrator

- Estimate direct illumination at x (as before)
- If volume: continue straight ahead until no volume (yields intersections y, z)



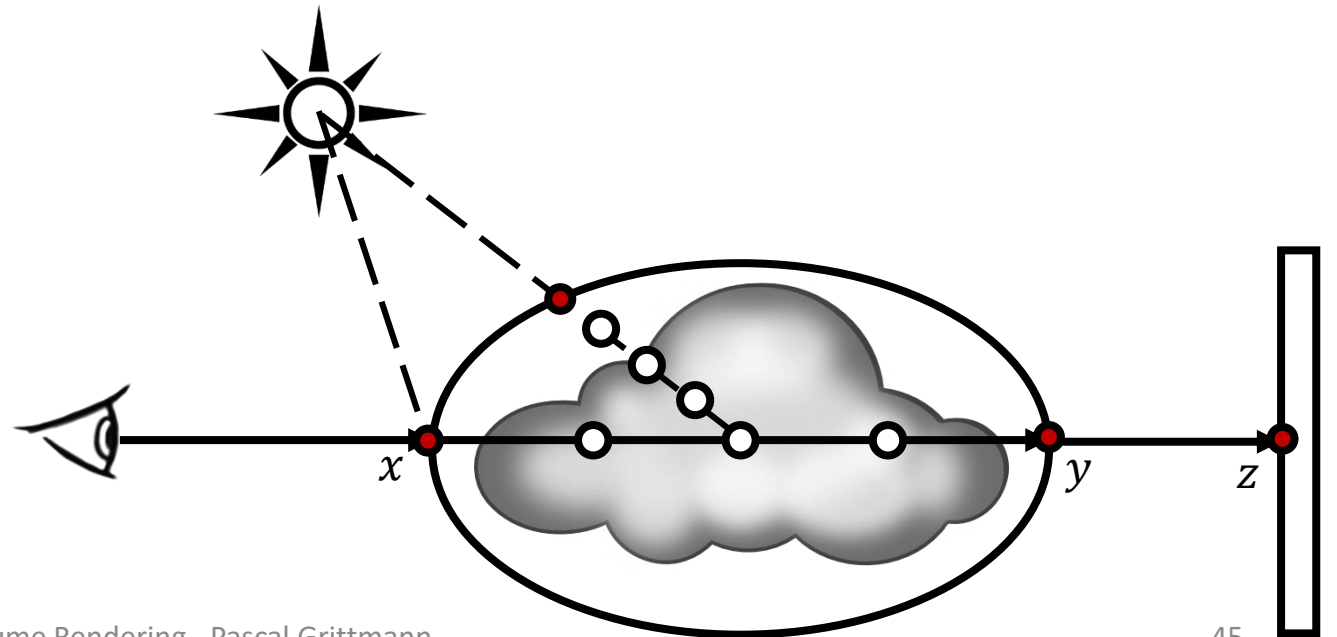
A Simple Volume Integrator

- Estimate direct illumination at x (as before)
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- Ray marching to estimate attenuation, emission,



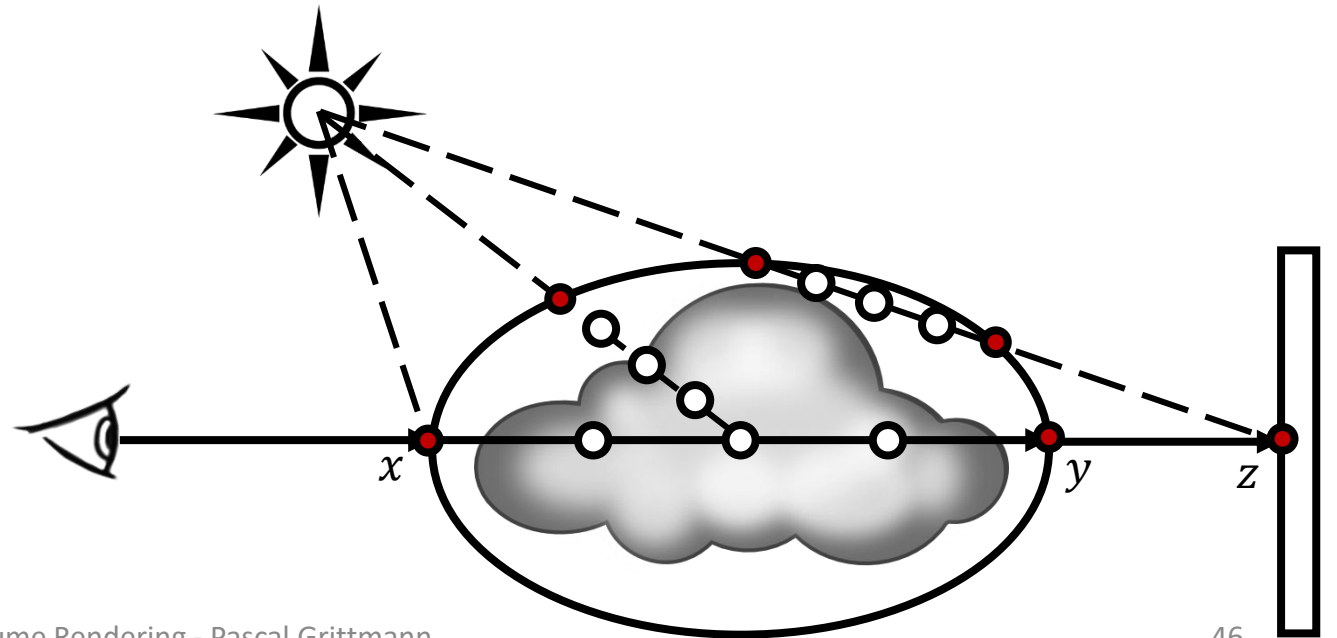
A Simple Volume Integrator

- Estimate direct illumination at x (as before)
- If volume: continue straight ahead until no volume (yields intersections y, z)
- Ray marching to estimate attenuation, emission, and in-scattering
 - Shadow rays to the lights + ray marching to compute attenuation



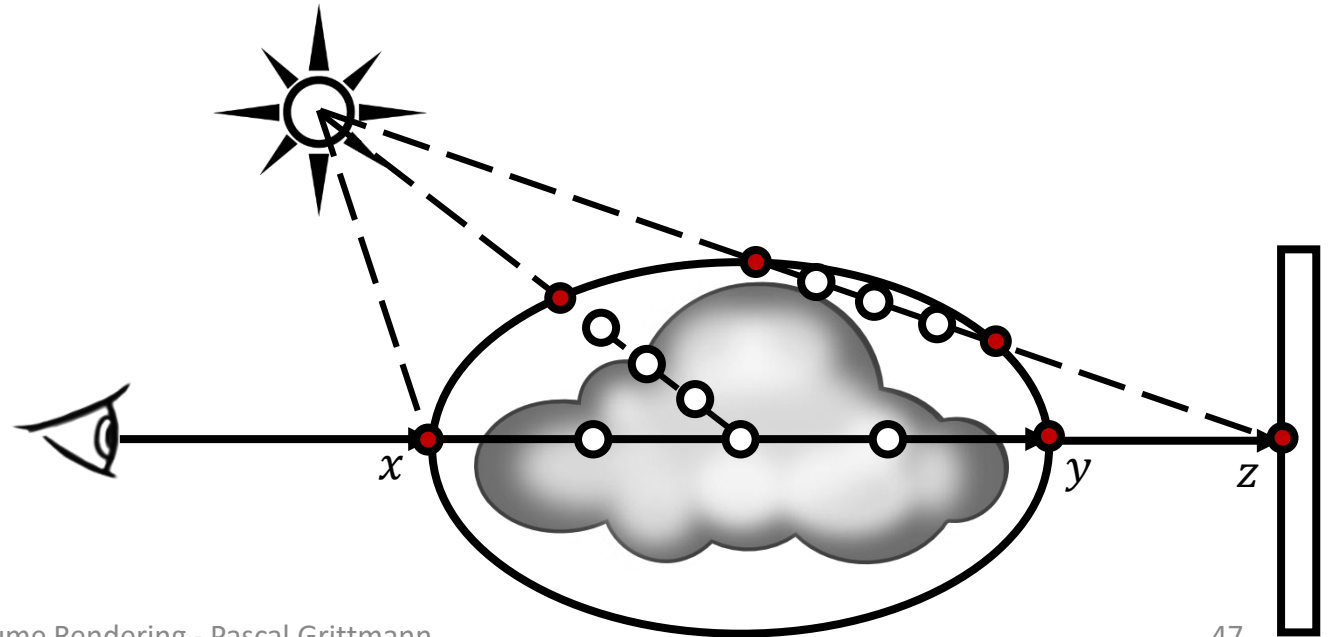
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- Compute illumination at z (as before)



A Simple Volume Integrator

- Estimate direct illumination at x (as before)
- If volume: continue straight ahead until no volume (yields intersections y, z)
- Ray marching to estimate attenuation, emission, and in-scattering
 - Shadow rays to the lights + ray marching to compute attenuation
- Compute illumination at z (as before)
- Add together:
 - Attenuated illumination from z
 - Volumetric emission along \overline{xy}
 - In-scattering along \overline{xy}
 - Direct illumination at x



A Simple Volume Integrator

- Estimate direct illumination at x (as before)
- If volume: continue straight ahead until no volume (yields intersections y, z)
- Ray marching to estimate attenuation, emission, and in-scattering
 - Shadow rays to the lights + ray marching to compute attenuation
- Compute illumination at z (as before)
- Add together:

- Multiply by BTDF!
- Attenuated illumination from z
 - Volumetric emission along \overline{xy}
 - In-scattering along \overline{xy}
 - Direct illumination at x

